

## 1 Properties of Nuclei

**nuclei:** core of the atom; consists of protons and neutrons.

**atomic number:**  $Z$ ; the number of protons in a nucleus.

**neutron number:**  $N$ ; the number of neutrons in a nucleus.

**nucleon number or mass number:**  $A = Z + N$ ; the number of nucleons in the nucleus.

### *Radii of nuclei*

The radii of nuclei are given approximately by the formula

$$R = R_0 A^{1/3},$$

where  $R_0$  is an empirical constant:

$$R_0 = 1.2 \times 10^{-15} \text{ m} = 1.2 \text{ fm.}$$

All nuclei have approximately the same density.

**nuclide:** a single nuclear species having specific values of both  $Z$  and  $N$ .

**isotopes:** nuclei of a given element with different mass numbers; different isotopes have slightly different physical properties.

**mass of proton:**  $m_p = 1.672621777 \times 10^{-27} \text{ kg}$

**mass of neutron:**  $m_n = 1.674927351 \times 10^{-27} \text{ kg}$

**unified atomic mass unit:**  $1 \text{ u} = 1.660538921 \times 10^{-27} \text{ kg}$ ; one-twelfth of the mass of the neutral carbon atom with mass number  $A = 12$ .

### Mass of $p$ , $n$ , $e$ in u

$$m_p = 1.007276466812 \text{ u}$$

$$m_n = 1.00866491600 \text{ u}$$

$$m_e = 5.4857990946 \times 10^{-4} \text{ u}$$

### Energy equivalent of 1 u from $E = mc^2$

$$E = (1.66054 \times 10^{-27} \text{ kg})(2.99792 \times 10^8 \text{ m/s})^2 = 1.49242 \times 10^{-10} \text{ J} = 931.494 \text{ MeV}$$

**binding energy:** magnitude of the total potential energy of the nucleus.

### *Mass defect*

The mass defect  $\Delta M$  for a nucleus with mass  $M$  containing  $Z$  protons and  $N$  neutrons is defined as

$$\Delta M = Zm_p + Nm_n - M.$$

*Mass defect is the mass difference. The total mass of a nucleus is always less than the total mass of its constituent parts due to the **mass equivalent** ( $E = mc^2$ ).*

**deuterium:** the isotope of hydrogen with mass number 2. The nucleus consists of a proton and a neutron bound together called the *deuteron*.

**nuclear spin:** the total angular momentum of a nucleus; generally associated with both orbital and spin angular momentum of the protons and neutrons.

## 2 Nuclear Stability

*nuclear force or strong force:*

- does not depend on charge; same for  $p$  and  $n$ .
- has a short range on the order of nuclear dimensions  $\sim 10^{-15}$  m.
- a particular nucleon interacts only with those few nucleons in its immediate vicinity (versus all other nucleons).
- favors the binding of pairs of  $p$  or  $n$  with opposite spins ( $p \uparrow$ - $p \downarrow$  or  $n \uparrow$ - $n \downarrow$ ) and of pairs of pairs ( $p$ - $p$  with  $n$ - $n$ ).

## 3 Radioactivity

**radioactivity:** the spontaneous disintegration of nuclides that are not stable.

**alpha decay:** decay of a nucleus emitting alpha ( $\alpha$ ) particles

**alpha particle:** identical to the  ${}^4\text{He}$  nucleus: 2  $p$  and 2  $n$  bound together; 0 total spin.

**beta-minus ( $\beta^-$ ) particle:** a free  $n$  that decays into a  $p$  and an  $e$  with an average lifetime of about 15 minutes.

**neutrino:**  $\nu$ ; note: the book has  $\nu_e$ , but there exist  $\nu_e$ ,  $\nu_\tau$ ,  $\nu_\mu$  neutrinos and all their anti-particles.

**anti-neutrino:**  $\bar{\nu}$ ;  $\bar{\nu}_e$ ,  $\bar{\nu}_\tau$ ,  $\bar{\nu}_\mu$

**gamma:**  $\gamma$ , as in gamma radiation can be emitted from an excited nucleus after an emission of an  $\alpha$  or a  $\beta^-$  particle.

### *Rate of decay of radioactive nuclei*

$$\frac{\Delta N}{\Delta t} = -\lambda N.$$

The constant  $\lambda$  is called the decay constant. It has different values for different nuclides.

### *Number of radioactive nuclei remaining after time $t$*

If there are  $N_0$  nuclei with decay constant  $\lambda$  at time  $t = 0$ , the number  $N$  remaining after time  $t$  is

$$N = N_0 e^{-\lambda t}.$$

**half-life:**  $T_{1/2}$  of a radioactive substance is the time required for the number of radioactive nuclei to decrease to half the original number  $N_0$ .

### *Decay constant and half-life*

$$T_{1/2} = \frac{\ln 2}{\lambda} = \frac{0.693}{\lambda}.$$

In particle physics, the life of an unstable nucleus or particle is usually described by the lifetime, not the half-life. The **lifetime** (or mean lifetime)  $T_{mean}$  is the average time for a nucleus or particle to decay:

$$T_{mean} = \frac{1}{\lambda} = \frac{T_{1/2}}{\ln 2} = \frac{T_{1/2}}{0.693}.$$

**activity:** of a specimen is the number of decays per unit time.

**curie:** Ci, common unit of activity; 1 Ci =  $3.70 \times 10^{10}$  decays/s.

**becquerel:** Bq, SI unit of activity; 1 Bq = 1 decay/s; and 1 Ci =  $3.70 \times 10^{10}$  Bq.

**radiocarbon dating:** the dating of archeological and geological specimens by measuring the concentration of radioactive isotopes.

## 4 Radiation and the Life Sciences

**radiation:** includes radioactivity ( $\alpha$ ,  $\beta$ ,  $\gamma$ ,  $n$ ) and electromagnetic radiation such as x rays and gamma rays. As these particles pass through matter, they lose energy, breaking molecular bonds and creating ions—*ionizing radiation*.

**radiation dosimetry:** quantitative description of the effect of radiation on living tissue.

**absorbed dose:** the energy delivered to the tissue, per unit mass of tissue.

**gray:** Gy, SI unit of absorbed dose of radiation.

**rad:** common unit of absorbed radiation.

### Definition of unit of absorbed radiation dose

$$1 \text{ Gy} = 1 \text{ J/kg}, 1 \text{ rad} = 0.01 \text{ J/kg} = 0.01 \text{ Gy} = 1 \text{ cGy}$$

**relative biological effectiveness (RBE):** a numerical factor weighting the type (energy) of radiation dose with x rays of 200 keV have 1 RBE.

Radiation	RBE
x rays and $\gamma$ rays	1
$e$	1
$p$	5
$\alpha$ particles	20
heavy ions	20
slow $n$	5-20
	(energy dependent)

**Sievert:** Sv, SI unit of equivalent dose for humans.

### Units of equivalent dose

$$\text{equivalent dose (Sv)} = \text{RBE} \times \text{absorbed dose (Gy)}.$$

A more common unit, corresponding to the rad, for equivalent dose for humans is the **rem**:

$$\text{equivalent dose (rem)} = \text{RBE} \times \text{absorbed dose (rad)}.$$

thus,  $1 \text{ rem} = 0.01 \text{ Sv}$ .

## 5 Nuclear Reactions

**nuclear reaction:** nuclear particles are rearranged as a result of the bombardment of a nucleus by a particle, rather than through a spontaneous natural process; obey classical conservation principles for electric charge, momentum, angular momentum, and energy *plus* conservation of the total number of nucleons.

**reaction energy:** the difference between the rest masses before and after the reaction.  $Q = (M_A + M_B - M_C - M_D)c^2$  where reaction energy  $Q$ , particles A and B  $\rightarrow$  particles C and D.

**exoergic reaction or exothermic reaction:**  $Q > 0$ , the total mass decreases and the total kinetic energy increases.

**endoergic or endothermic:**  $Q < 0$ , the total mass increases and the total kinetic energy decreases.

**threshold energy:** the energy required for an endoergic reaction:  $E \geq |Q|$ .

## 6 Nuclear Fission

*nuclear fission*: a decay process in which an unstable nucleus splits into two fragments of comparable mass instead of emitting an  $\alpha$  or  $\beta$  particle.

*fission fragments*: products of fission.

*induced fission*: fission resulting from neutron absorption.

*spontaneous fission*: fission occurring without initial neutron absorption.

*chain reaction*: fission that induces more fission events.

*nuclear reactor*: a system in which a controlled nuclear chain reaction is used to liberate energy.

## 7 Nuclear Fusion

*nuclear fusion*: two or more small light nuclei combine to form a larger nucleus.

*thermonuclear reactions*: fusion chain reactions

*cold fusion*: fusion not requiring high temperatures; hypothetical.

## 8 Fundamental Particles

*positron*:  $\beta^+$  or  $e^+$

*anti-particles*: particle with same mass but charge is equal and opposite of partner.

*pair production*: products of high-energy collisions of charged particles or gamma rays with matter.

*muons*:  $\mu^-$  and its antiparticle  $\mu^+$ ; charge equal to electron; mass  $207m_e$ .

*pions*:  $\pi^\pm$ ; mass  $273m_e$ ; charge  $e^\pm$

## 9 High-Energy Physics

### Four Forces

- the strong interaction: gluons
- the electromagnetic interaction: photons
- the weak interaction: W & Z bosons
- the gravitational interaction: gravitons

*antiproton*:  $\bar{p}$

*hadrons*: have strong interactions; include baryons and mesons;

*leptons*: do not have strong interactions;  $e^\pm, \mu^\pm, \tau^\pm, \nu_{e,\tau,\mu}$ ; obey conservation principle

*bosons*: zero or integer spins;

*fermions*: half-integer spins;

*baryons*: fermions; nucleons and more massive particles that resemble nucleons; hyperons:  $\Lambda, \Sigma, \Xi, \Omega$ ;

*mesons*: spin 0 or 1; boson;  $\pi^\pm$ , K mesons or kaons,  $\eta$  mesons;

*conservation of baryon number*: all baryons have a baryon number of  $\pm 1$  and in all interactions, the total baryon number is conserved.

*strangeness*: a quantity or property of mesons; conserved in production processes, but not in individual decay processes.

*absolute conservation laws*: always true such as baryon number conservation.

*conditional conservation law*: conservation laws that are conditional such as strangeness.

*quarks*: fundamental particles (for the time being) that make up baryons (3), mesons

*gluons*: force mediators involved in strong force; massless spin-1 bosons.

*quantum chromodynamics*: theory of the strong interactions.

*standard model*: the current scheme of 6 quarks, 6 leptons, and the force mediators and our understanding of their behavior.

*electroweak theory*: theory that treats the weak and electromagnetic forces as two aspects of a single interaction at sufficiently high energies.

*grand unified theories or GUTs*: comprehensive theory of strong, weak, and electromagnetic interactions.

*theory of everything*: GUTs plus gravity.

## 10 Cosmology

*red shifts*: longer wavelengths of distant galaxies that indicate expansion of the universe.

*Hubble's law*: the speed of recession,  $v$ , of a galaxy is approximately proportional to its distance  $r$  from us.

*Big Bang*: the hypothesis, suggested by Hubble's law, that at some time in the past, all the matter in the universe was concentrated in a very small space and was blown apart in an immense explosion—the Big Bang, giving all observable matter more or less the velocities observed today.

*dark matter*: about 26.8% of energy in the universe of unknown content. CMB

*dark energy*: about 69% of the energy in the universe. universe. CMB

These values for *dark energy* and *dark matter* are different from your text. This happens in REAL science...values change as we know more and theories are refined! Keep your mind open, keep questioning, keep thinking.

## 11 Links

Particle Poster .jpg

### **Pdfs**

Fermions

Bosons

Baryons

Mesons

Atomic Structure

Properties of Interactions

Particle Processes

Unsolved Mysteries

History of the Universe Images

Particle Overview Simple

### **Awesome Sites**

The Particle Adventure

The Official String Theory Web Site

The Particle Zoo